

BLADE

Allen Telescope Array GPU accelerated beamformer backend

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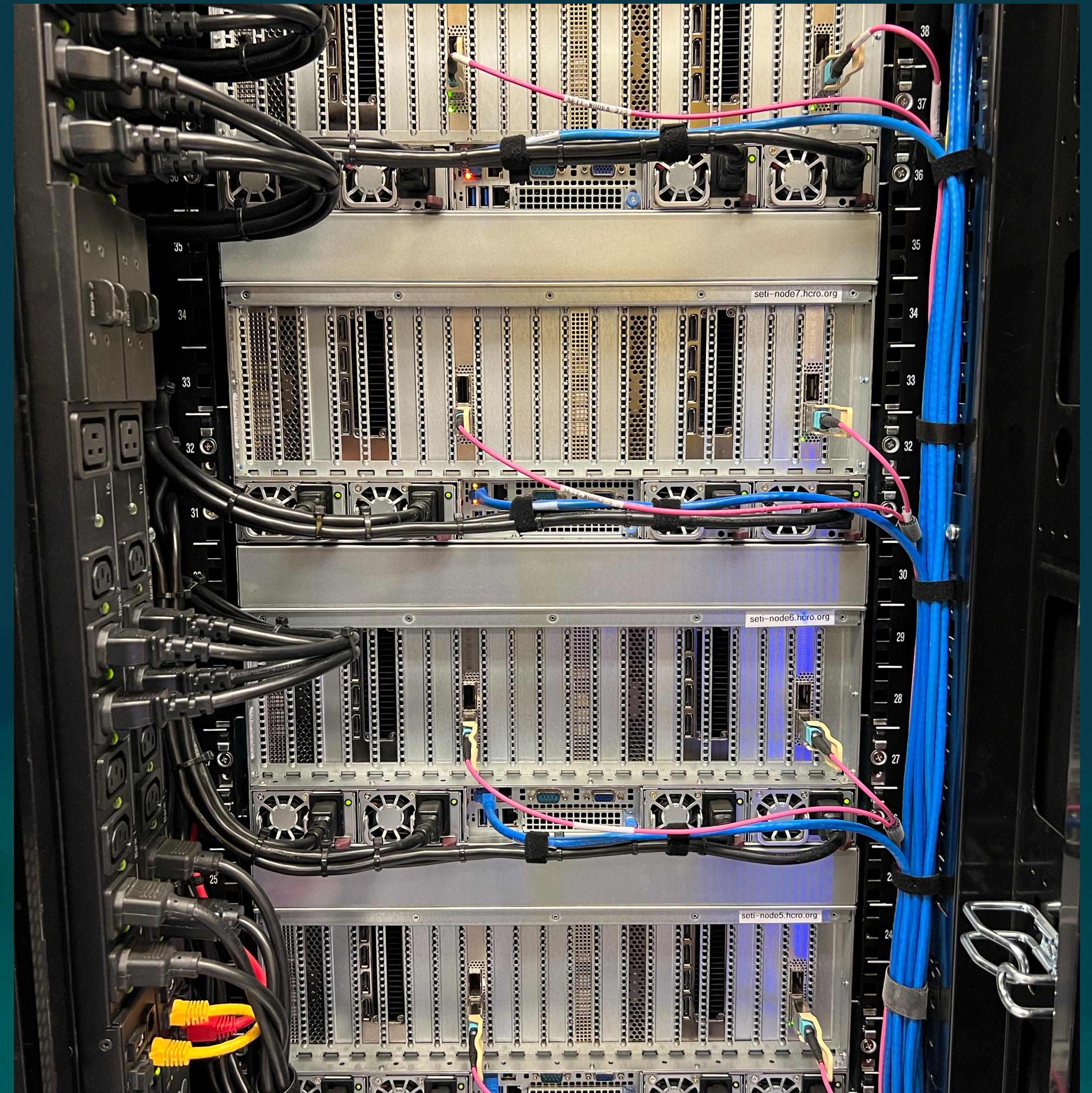
The Equipment

- Designed primarily for SETI purposes.
- Composed of 42 antennas each with 6.1 meters in diameter.
- Randomly scattered and have a maximum baseline of 300 meters.
- Sensitive to a wide and continuous range of frequencies ranging from 900 MHz to 12 GHz.
- Maximum bandwidth is around 1.5 GHz per antenna for each polarization.
- The GPU compute cluster receive digitized data via a 100 Gbps network.
- More details will be covered by a presentation from our Chief Astronomer, Wael Farah, on Thursday (Session 451).



Digital Signal Processing

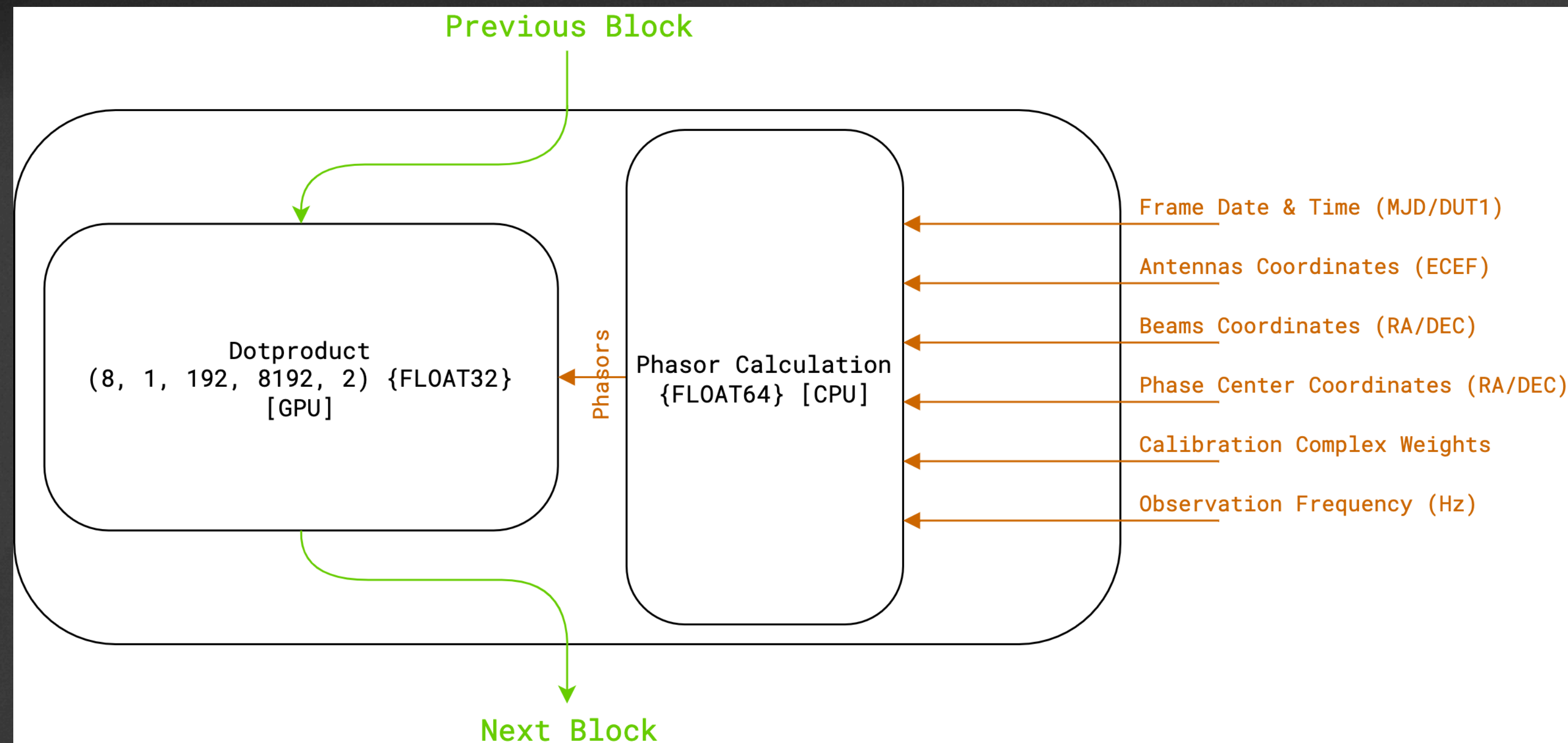
- Always looking at ways to expand its capabilities.
- Development cycles where the target is set, the code and tests are written, and the new pipeline is tested on sky.
- Brand-new processing framework called BLADE the “Breakthrough Listen Accelerated DSP Engine”.
- Compute cluster composed of 8 dual-socket EPYC servers with two NVIDIA RTX 3090 class graphics cards.
- Processing performed by the GPU.
- Composed of modules inside pipelines organized with runners.
- Modular design allows development without affecting the stability of production sections.



Beamformer

BLADE Processing Pipeline

- First to be developed.
- Combines the signal received by individual antennas into a single larger disk.
- Accomplished by delaying the signal slightly using a phasor.
- Supports electronic steering and multiple beam generation.



High-Resolution Spectrogram

BLADE Processing Pipeline

- Received data is pre-channelized down to 500 kHz in resolution.
- The ideal channelization resolution is around 1 Hz for SETI search.
- Module further channelizes the data using FFT.
- Total of 1.5 million bins FFT of a time-domain vector for each polarization.
- Translates into a 3 million bins (3 GHz) FFT done every second.

Detection

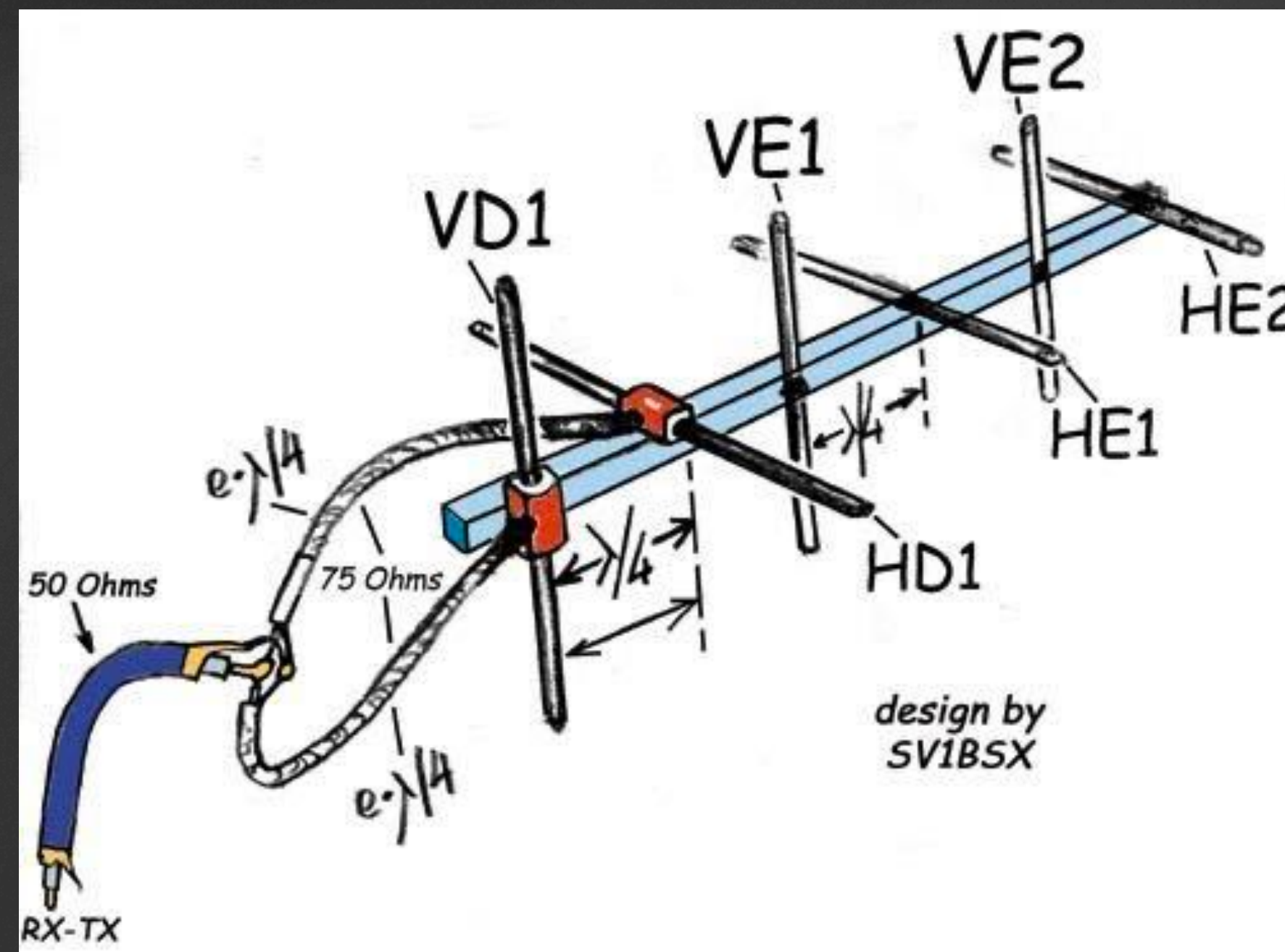
BLADE Processing Pipeline

- Calculates the Stokes-I.
- Integrates the signal in real-time.
- Integration does not require writing data to disk.
- Continuous operation until overflow.

Circular Polarization

BLADE Processing Pipeline

- Data produced by the telescope have Vertical and Horizontal polarizations.
- Generates the RHCP and LHCP.
- Necessary calibration done internally by the pipeline.
- Useful for telecom purposes and some science.



Future Modules

BLADE Processing Pipeline

- Coherent de-dispersion (pulsars and drift-search).
- Active RFI mitigation using the beamformer.
- Correlation.

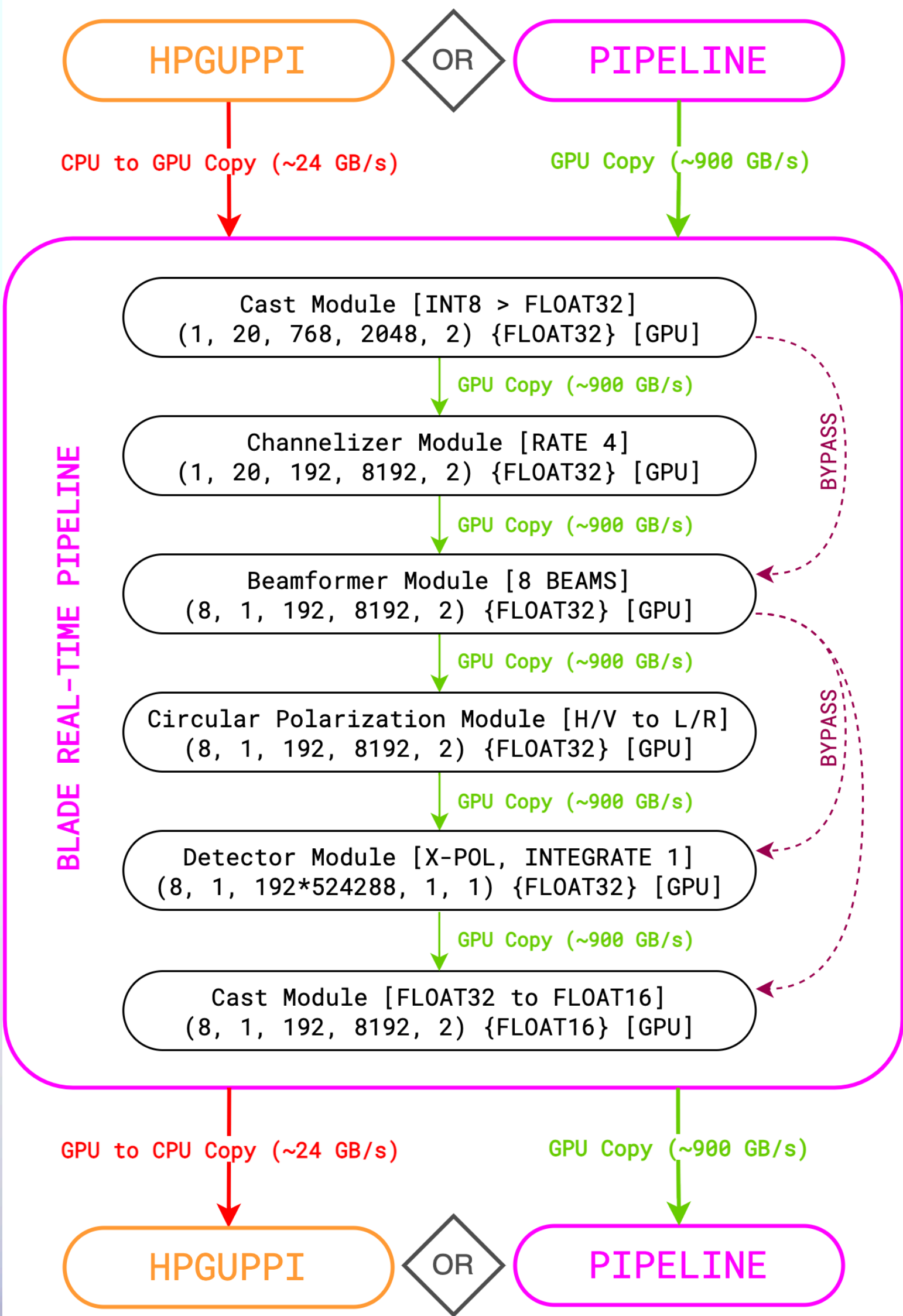
Takeaways

- The Allen Telescope Array is a good example of an instrument primarily created for SETI being used for other scientific purposes.
- Fundamental to design the instrument to accommodate incremental changes without impacting on-sky time.
- Real-time processing pipelines increase efficiency.

Pipelines

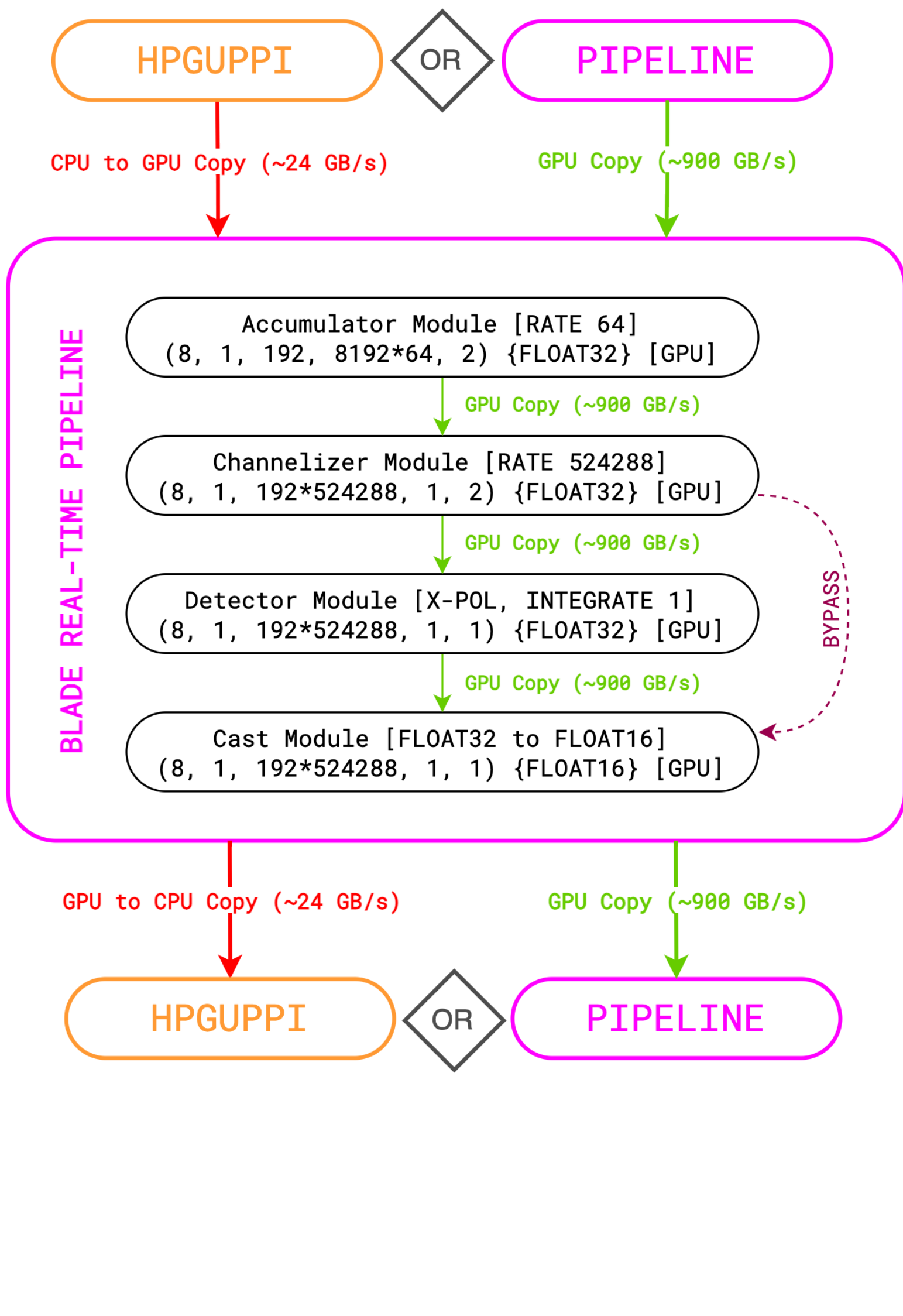
ATA PIPELINE - MODE-B

Beamforming/Detection



GENERIC PIPELINE - MODE-H

High Resolution Channelization (<1Hz/bin resolution)



Thanks for listening!

<https://github.com/luigifcruz/blade>

Questions?

Contact me!

<https://luigi.ltd/contact/>

